

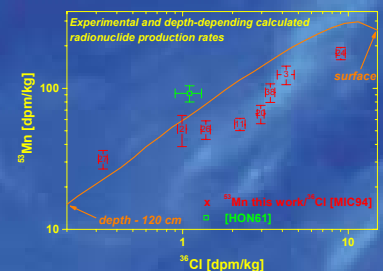
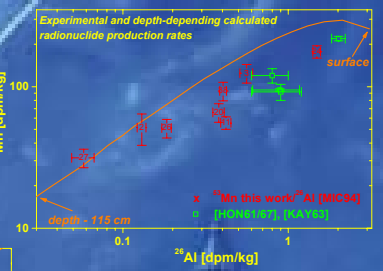
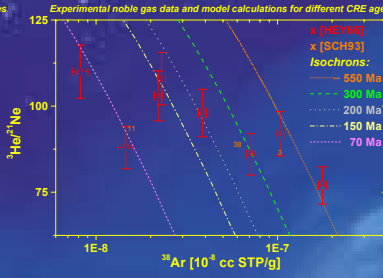
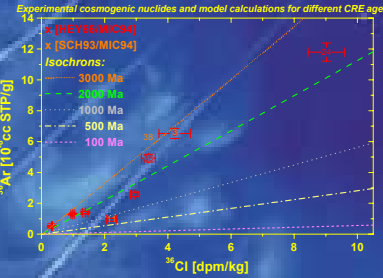
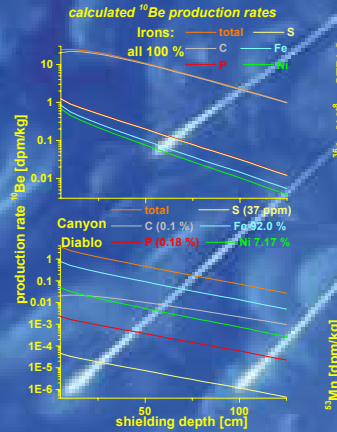
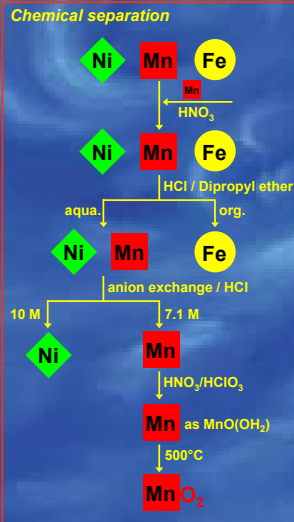
# COSMOGENIC NUCLIDES IN IRON METEORITES: CHALLENGING CANYON DIABLO

Silke Merchel, T. Faestermann, U. Herpers, K. Knie, G. Korschinek, I. Leya, M. E. Lipschutz, R. Michel



## Project:

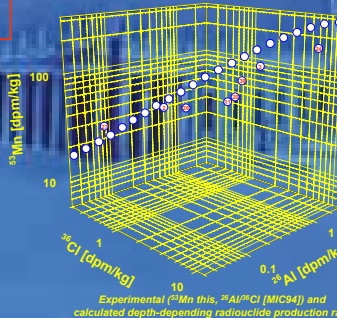
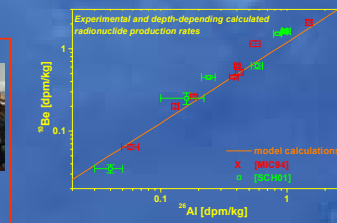
- 8 samples from a suite of 56 Canyon Diablo fragments from known locations
- well-characterized samples: stable and lighter radionuclides already published [HEY66, MIC94]
- here: first AMS-<sup>53</sup>Mn depth-profile of an iron meteorite (first stony, see [MER97])



### <sup>53</sup>Mn AMS measurements @ Munich

(for details see [1000])

sample form	Mn(OH) <sub>2</sub>
mixed with	-
sample holder	Cu
extraction as	MnO
charge state	11+
ion energy	160 MeV
measured ratios	7E-10 - 4E-9
background	3E-14
transmission	~3 %
standard	<sup>51</sup> V(c,2n) <sup>53</sup> Mn



### Data for model calculations

(for details see [LEY09a])

J <sub>0</sub>	4.36 cm <sup>-2</sup> s <sup>-1</sup> MeV <sup>-1</sup>
M	550 MeV
σ	[MER00, MIC97]
Φ	[MAS99, SCH99]
chemical comp.	[BUC75, MET99]

### Results of AMS measurements for Canyon Diablo samples

Sample <sup>a</sup>	II	III	11	20	24	26	27	38
<sup>53</sup> Mn/ <sup>55</sup> Mn [ε-9]	0.65 ± 0.16	2.60 ± 0.39	1.31 ± 0.13	1.50 ± 0.23	4.10 ± 0.41	1.07 ± 0.16	0.69 ± 0.10	2.31 ± 0.35
<sup>53</sup> Mn activity [dpm/kg]	51 ± 13	124 ± 19	55.1 ± 5.5	65.4 ± 9.8	175 ± 18	50.6 ± 7.6	31.2 ± 4.7	93 ± 14
<sup>53</sup> Mn production rate <sup>a</sup> [atoms·min <sup>-1</sup> ·kg <sup>-1</sup> ]	51 ± 13	125 ± 19	55.7 ± 5.6	66.0 ± 9.9	177 ± 18	51.1 ± 7.7	31.5 ± 4.7	94 ± 14

<sup>a</sup> Sample numbers used by [HEY66], II and III refer to the Canyon Diablo 2 and 3 meteorites respectively, corrected for terrestrial age of 64 ka [BUC75]

### Conclusions:

- Because of the **complex exposure** history of Canyon Diablo
- ⇒ the commonly used combination of radionuclide/stable noble gas nuclide should **NOT** be used to determine neither CRE ages nor terrestrial ages!
- ⇒ noble gas ratios like <sup>3</sup>He/<sup>21</sup>Ne or <sup>4</sup>He/<sup>21</sup>Ne should **NOT** be used as shielding indicators!
- lighter radionuclides should be used with **utmost precaution** due to strong dependence of chemical inhomogeneity as shielding or terrestrial age indicator!
- combination of heavier radionuclides, e.g. <sup>41</sup>Ca, <sup>53</sup>Mn, <sup>50</sup>Ni, <sup>60</sup>Fe most reliable as shielding or terrestrial age indicator!
- Canyon Diablo is an ideal object to test the quality of model calculations of radionuclides in iron meteorites, but difficult for noble gases!
- Improvement on model calculations is still necessary (especially for heavier nuclides)!**

### Acknowledgements:

This work was partially funded by the DFG. We gratefully acknowledge partially spectra data by J. Masarik [MAS99] and unpublished noble gas data from L. Schultz [SCH93].

**References:**

[BUC75] V. F. Buchwald, Handbook of Iron Meteorites, University of California Press, Berkeley, Los Angeles, London (1975).

[HEY66] D. Heymann et al., Canyon Diablo Ironites: Radiographic and Mass Spectrometric Study of 56 Fragments, *J. Geophys. Res.* 71 (1966) 619.

[HON61] M. Honda et al., Radioactive species produced by cosmic rays in iron meteorites, *Geochim. Cosmochim. Acta* 22 (1961) 135-154 (from [NS87]).

[HON67] M. Honda and Arnold J. R. (1967) Effects of cosmic rays on meteorites, *Nature*, *Physik* 462 (1967) pp. 813, Springer-Verlag, Berlin-Heidelberg (from [NS87]).

[KOR93] G. Korschinek, Cosmic-ray induced radionuclides in iron meteorites, *Geochim. Cosmochim. Acta* 57 (1993) pp. 427-437 (from [NS87]).

[KNE00] K. Knie et al., High-sensitivity AMS for heavy nuclides at the Munich Tandem accelerator, *Nucl. Instr. and Meth. in Phys. Res. B* 172 (2000) 717-720.

[LEY09a] I. Leya et al., The production of cosmogenic nuclides in stony meteoroids by galactic cosmic ray particles, *Meteor. Planet. Sci.* 35 (2000) 299-326.

[LEY09b] I. Leya et al., Simulation of the interaction of GCR protons with meteoroids - On the Production of Radionuclides in Thick Targets and Iron Targets irradiated isotropically with 1.6 GeV Protons, *Meteorit. Planet. Sci.* 35 (2000) 287-316.

[MAS99] J. Masarik, *priv. com.* (1999).

[MER97] S. Merchel, et al., DEPTH PROFILES OF LONG-LIVED COSMOGENIC RADIONUCLIDES IN IRON, *Meteorit. Planet. Sci.* 32 (1997) A96.

[MIC94] S. Michel et al., Iron and back-target cross sections for the production of <sup>53</sup>Mn and <sup>56</sup>Fe, *Nucl. Instr. and Meth. in Phys. Res. B* 112 (2000) 806-811.

[NS87] M. Honda, Version 4.0 for Windows 1.0, Meteorite Data Retrieval Software, Cyber Hobby.

[MIC94] E. S. Michlovich et al., Aluminum 26, <sup>10</sup>Be, and <sup>14</sup>C depth profiles in the Canyon Diablo iron meteorite, *J. Geophys. Res.* 99 (1994) 23157-23164.

[SCH93] R. Michel et al., Cross sections for the production of residual nuclides by low- and medium-energy protons from the target elements C, N, O, Mg, Al, Si, Ca, Ti, V, Mn, Fe, Co, Ni, Cu, Sr, Y, Zr, Nb, Ba and Au, *Nucl. Instr. and Meth. in Phys. Res. B* 120 (1997) 185-193.

[NS87] K. Nishikubo, <sup>10</sup>Be, <sup>14</sup>C, and <sup>14</sup>C in METEORITES: DATA COMPILATION, *Nucl. Tracks Radiat. Meas.* 13 (1987) 209-213.

[SCH93] L. Schultz, *priv. com.* (1993).

[SCH99] C. Schimmel et al., Shock Melting of the Canyon Diablo Impactor: Constraints from Nickel-60 Contents and Numerical Modeling, *Science* 285 (1999) 85-89.

[SCH99] C. Schimmel et al., TERRESTRIAL AGES OF CANYON DIABLO METEORITES, *Meteorit. Planet. Sci.* 34 (2001) 883-892, this conference.

[SUT78] S. R. Sutton, Thermoluminescence measurements on shock-metamorphosed sanidine and dolomite from Meteor Crater, Arizona, 2. Thermoluminescence age of Meteor Crater, *J. Geophys. Res.* 80 (1985) 3696-3700.